NuFIT 3.2: Three-neutrino fit based on data available in January 2018

Ivan Esteban,^{*a*} M. C. Gonzalez-Garcia,^{*a,b,c*} Alvaro Hernandez-Cabezudo,^{*d*} Michele Maltoni,^{*e*} Ivan Martinez-Soler,^{*e*} Thomas Schwetz^{*d*}

- ^aDepartament de Fisíca Quàntica i Astrofísica and Institut de Ciencies del Cosmos, Universitat de Barcelona, Diagonal 647, E-08028 Barcelona, Spain
- ^bInstitució Catalana de Recerca i Estudis Avançats (ICREA), Pg. Lluis Companys 23, 08010 Barcelona, Spain.
- ^cC.N. Yang Institute for Theoretical Physics, State University of New York at Stony Brook, Stony Brook, NY 11794-3840, USA
- ^dInstitut für Kernphysik, Karlsruher Institut für Technologie (KIT), D-76021 Karlsruhe, Germany

^eInstituto de Física Teórica UAM/CSIC, Calle de Nicolás Cabrera 13–15, Universidad Autónoma de Madrid, Cantoblanco, E-28049 Madrid, Spain

E-mail: ivan.esteban@fqa.ub.edu,

maria.gonzalez-garcia@stonybrook.edu, alvaro.cabezudo@kit.edu, michele.maltoni@csic.es, ivanj.m@csic.es, schwetz@kit.edu

ABSTRACT: We present updated results for our global analysis of solar, atmospheric, reactor, and accelerator neutrino data in the framework of three-neutrino oscillations. We also provide χ^2 tables for the various one- and two-dimensional projections of the global analysis. If you use these results, please refer to both [1] and [2]. Data sets which have been updated with respect to NuFIT 3.1 are marked by the " \Rightarrow " tag.

Solar experiments

- External information: Standard Solar Model [3].
- Chlorine total rate [4], 1 data point.
- Gallex & GNO total rates [5], 2 data points.
- SAGE total rate [6], 1 data point.
- SK1 full energy and zenith spectrum [7], 44 data points.
- SK2 full energy and day/night spectrum [8], 33 data points.
- SK3 full energy and day/night spectrum [9], 42 data points.
- SK4 2055-day day-night asymmetry [10] and 2365-day energy spectrum [11], 24 data points.
- SNO combined analysis [12], 7 data points.
- Borexino Phase-I 741-day low-energy data [13], 33 data points.
- Borexino Phase-I 246-day high-energy data [14], 6 data points.
- Borexino Phase-II 408-day low-energy data [15], 42 data points.

Atmospheric experiments

- External information: Atmospheric neutrino fluxes [16].
- IceCube/DeepCore 3-year data [17, 18], 64 data points.

Reactor experiments

- KamLAND separate DS1, DS2, DS3 spectra [19] with Daya-Bay reactor ν fluxes [20], 69 data points.
- Double-Chooz FD-I/ND and FD-II/ND spectral ratios, with 455-day (FD-I), 363-day (FD-II) and 258-day (ND) exposures [21], 56 data points.
- Daya-Bay 1230-day EH2/EH1 and EH3/EH1 spectral ratios [22], 70 data points.
- Reno 1500-day FD/ND spectral ratios [23], 26 data points.

Accelerator experiments

- MINOS 10.71×10^{20} pot ν_{μ} -disappearance data [24], 39 data points.
- MINOS 3.36×10^{20} pot $\bar{\nu}_{\mu}$ -disappearance data [24], 14 data points.
- MINOS 10.6×10^{20} pot ν_e -appearance data [25], 5 data points.
- MINOS 3.3×10^{20} pot $\bar{\nu}_e$ -appearance data [25], 5 data points.
- T2K 14.93×10^{20} pot ν_{μ} -disappearance data [26], 55 data points.
- T2K 14.93×10^{20} pot ν_e -appearance data [26], 39 data points.
- T2K 7.62 × 10²⁰ pot $\bar{\nu}_{\mu}$ -disappearance data [26], 55 data points.
- T2K 7.62 × 10²⁰ pot $\bar{\nu}_e$ -appearance data [26], 23 data points.
- \Rightarrow NO ν A 8.85 \times 10²⁰ pot ν_{μ} -disappearance data [27], 72 data points.
- \Rightarrow NO ν A 8.85 \times 10²⁰ pot ν_e -appearance data [27], 19 data points.

Description of the χ^2 data tables

We provide two gzip-compressed files (one for Normal and one for Inverted Ordering) containing the χ^2 data tables for our global analysis. Each file is divided into 21 sections, identified by a unique tag and corresponding to a particular one- or two-dimensional projections. The tags and the meaning of the data columns for each section are listed below (note that $\ell = 1$ for NO and $\ell = 2$ for IO).

N°	Section tag	1^{st} column	2 nd column	3 rd column
1	# T13/T12	$\sin^2 \theta_{13}$	$\sin^2 \theta_{12}$	$\Delta \chi^2$
2	# T13/DMS	$\sin^2 \theta_{13}$	$\log_{10} \left(\Delta m_{21}^2 / [\mathrm{eV}^2] \right)$	$\Delta \chi^2$
3	# T12/DMS	$\sin^2 \theta_{12}$	$\log_{10} \left(\Delta m_{21}^2 / [\mathrm{eV}^2] \right)$	$\Delta\chi^2$
4	# T13/T23	$\sin^2 \theta_{13}$	$\sin^2 \theta_{23}$	$\Delta \chi^2$
5	# T13/DMA	$\sin^2 \theta_{13}$	$\Delta m_{3\ell}^2 / [10^{-3} \mathrm{eV^2}]$	$\Delta \chi^2$
6	# T23/DMA	$\sin^2 \theta_{23}$	$\Delta m_{3\ell}^2 / [10^{-3} \text{ eV}^2]$	$\Delta \chi^2$

N°	Section tag	1 st column	2 nd column	3 rd column
7	# T13/DCP	$\sin^2 \theta_{13}$	$\delta_{ m CP} / [m deg]$	$\Delta \chi^2$
8	# T23/DCP	$\sin^2 \theta_{23}$	$\delta_{ m CP} / [m deg]$	$\Delta \chi^2$
9	# DMA/DCP	$\Delta m_{3\ell}^2 / [10^{-3} \mathrm{eV^2}]$	$\delta_{ m CP} / [m deg]$	$\Delta \chi^2$
10	# T12/T23	$\sin^2 \theta_{12}$	$\sin^2 heta_{23}$	$\Delta \chi^2$
11	# T12/DCP	$\sin^2 \theta_{12}$	$\delta_{ m CP} / [m deg]$	$\Delta \chi^2$
12	# T12/DMA	$\sin^2 \theta_{12}$	$\Delta m_{3\ell}^2 / [10^{-3} \text{ eV}^2]$	$\Delta\chi^2$
13	# DMS/T23	$\log_{10}\left(\Delta m_{21}^2 /[\mathrm{eV}^2]\right)$	$\sin^2 heta_{23}$	$\Delta \chi^2$
14	# DMS/DCP	$\log_{10}\left(\Delta m_{21}^2 / [\mathrm{eV}^2]\right)$	$\delta_{ m CP} / [m deg]$	$\Delta \chi^2$
15	# DMS/DMA	$\log_{10}\left(\Delta m_{21}^2 /[\mathrm{eV}^2]\right)$	$\Delta m_{3\ell}^2 / [10^{-3} \text{ eV}^2]$	$\Delta \chi^2$
16	# T13	$\sin^2 \theta_{13}$	$\Delta\chi^2$	
17	# T12	$\sin^2 \theta_{12}$	$\Delta \chi^2$	
18	# T23	$\sin^2 heta_{23}$	$\Delta \chi^2$	
19	# DCP	$\delta_{ m CP} / [m deg]$	$\Delta \chi^2$	
20	# DMS	$\int \log_{10} \left(\Delta m_{21}^2 / [\mathrm{eV}^2] \right)$	$\Delta \chi^2$	
21	# DMA	$\Delta m_{3\ell}^2 / [10^{-3} \text{ eV}^2]$	$\Delta \chi^2$	

References

- [1] I. Esteban, M. C. Gonzalez-Garcia, M. Maltoni, I. Martinez-Soler and T. Schwetz, Updated fit to three neutrino mixing: exploring the accelerator-reactor complementarity, 1611.01514.
- [2] I. Esteban, M. C. Gonzalez-Garcia, A. Hernandez-Cabezudo, M. Maltoni, I. Martinez-Soler and T. Schwetz, "NuFIT 3.2 (2018)." http://www.nu-fit.org.
- [3] N. Vinyoles, A. M. Serenelli, F. L. Villante, S. Basu, J. Bergström, M. C. Gonzalez-Garcia et al., A new Generation of Standard Solar Models, Astrophys. J. 835 (2017) 202, [1611.09867].
- [4] B. T. Cleveland et al., Measurement of the solar electron neutrino flux with the Homestake chlorine detector, Astrophys. J. 496 (1998) 505–526.
- [5] F. Kaether, W. Hampel, G. Heusser, J. Kiko and T. Kirsten, *Reanalysis of the GALLEX solar neutrino flux and source experiments*, *Phys. Lett.* B685 (2010) 47–54, [1001.2731].
- [6] SAGE collaboration, J. N. Abdurashitov et al., Measurement of the solar neutrino capture rate with gallium metal. III: Results for the 2002–2007 data-taking period, Phys. Rev. C80 (2009) 015807, [0901.2200].
- [7] SUPER-KAMIOKANDE collaboration, J. Hosaka et al., Solar neutrino measurements in Super-Kamiokande-I, Phys. Rev. D73 (2006) 112001, [hep-ex/0508053].
- [8] SUPER-KAMIOKANDE collaboration, J. Cravens et al., Solar neutrino measurements in Super-Kamiokande-II, Phys. Rev. D78 (2008) 032002, [0803.4312].
- [9] SUPER-KAMIOKANDE collaboration, K. Abe et al., Solar neutrino results in Super-Kamiokande-III, Phys. Rev. D83 (2011) 052010, [1010.0118].
- [10] Y. Nakano, ⁸B solar neutrino spectrum measurement using Super-Kamiokande IV. PhD thesis, Tokyo U., 2016-02.

- [11] Y. Nakano, "Solar neutrino results from Super-Kamiokande." Talk given at the 38th International Conference on High Energy Physics, Chicago, USA, August 3–10, 2016.
- [12] SNO collaboration, B. Aharmim et al., Combined Analysis of all Three Phases of Solar Neutrino Data from the Sudbury Neutrino Observatory, 1109.0763.
- [13] BOREXINO collaboration, G. Bellini et al., Precision measurement of the 7Be solar neutrino interaction rate in Borexino, Phys. Rev. Lett. 107 (2011) 141302, [1104.1816].
- BOREXINO collaboration, G. Bellini et al., Measurement of the solar 8B neutrino rate with a liquid scintillator target and 3 MeV energy threshold in the Borexino detector, Phys. Rev. D82 (2010) 033006, [0808.2868].
- [15] BOREXINO collaboration, G. Bellini et al., Neutrinos from the primary proton-proton fusion process in the Sun, Nature 512 (2014) 383–386.
- [16] M. Honda, M. Sajjad Athar, T. Kajita, K. Kasahara and S. Midorikawa, Atmospheric neutrino flux calculation using the NRLMSISE-00 atmospheric model, Phys. Rev. D92 (2015) 023004, [1502.03916].
- [17] ICECUBE collaboration, M. Aartsen et al., Determining neutrino oscillation parameters from atmospheric muon neutrino disappearance with three years of IceCube DeepCore data, Phys. Rev. D91 (2015) 072004, [1410.7227].
- [18] ICECUBE collaboration, J. P. Yañez et al., "IceCube Oscillations: 3 years muon neutrino disappearance data." http://icecube.wisc.edu/science/data/nu_osc.
- [19] KAMLAND collaboration, A. Gando et al., Reactor On-Off Antineutrino Measurement with KamLAND, Phys. Rev. D88 (2013) 033001, [1303.4667].
- [20] DAYA BAY collaboration, F. P. An et al., Improved Measurement of the Reactor Antineutrino Flux and Spectrum at Daya Bay, Chin. Phys. C41 (2017) 013002, [1607.05378].
- [21] A. Cabrera Serra, "Double Chooz Improved Multi-Detector Measurements." Talk given at the *CERN EP colloquium*, CERN, Switzerland, September 20, 2016.
- [22] DAYA BAY collaboration, F. P. An et al., Measurement of electron antineutrino oscillation based on 1230 days of operation of the Daya Bay experiment, Phys. Rev. D95 (2017) 072006, [1610.04802].
- [23] H. Seo, "New Results from RENO." Talk given at the EPS Conference on High Energy Physics, Venice, Italy, July 5–12, 2017.
- [24] MINOS collaboration, P. Adamson et al., Measurement of Neutrino and Antineutrino Oscillations Using Beam and Atmospheric Data in MINOS, Phys. Rev. Lett. 110 (2013) 251801, [1304.6335].
- [25] MINOS collaboration, P. Adamson et al., Electron neutrino and antineutrino appearance in the full MINOS data sample, Phys. Rev. Lett. (2013), [1301.4581].
- [26] A. Izmaylov, "T2K Neutrino Experiment. Recent Results and Plans." Talk given at the Flavour Physics Conference, Quy Nhon, Vietnam, August 13–19, 2017.
- [27] A. Radovic, "Latest oscillation results from NOvA." Joint Experimental-Theoretical Physics Seminar, Fermilab, USA, January 12, 2018.